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Harmonic potential as an effective limit of a discrete classical interaction

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Abstract

Motivated by improving the understanding of the quantum-to-classical transition we use a simple model of classical discrete interactions for studying the discrete-to-continuous transition in the classical harmonic oscillator. A parallel is traced with gravity to stress the relevance of such discrete interaction models.

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Fundamental interactions, according to quantum field theory, are realized through the exchange of interaction quanta-packets of matter-energy with defined quantum numbers, namely momentum-energy, spin, electric charge, etc. They are discrete interactions, in contradistinction to the classical continuous picture. The Bohr correspondence principle, a useful guideline in the early days of quantum mechanics, states that in the limit of very large quantum numbers the classical idea of continuity must result from the quantum discreteness as an effective description. It would be very interesting to see in a clear way that this discreteto-continuous transition occurs. This is the objective of the present paper with the use of a simple model of discrete classical interaction for studying this transition in the classical simple harmonic oscillator. We should not forget, however, that the harmonic potential, although an extremely useful tool in all branches of modern physics, is not itself a fundamental interaction, which, as well known, are just the gravitational, electromagnetic, weak and strong interactions; actually it is just an effective description. This may just valorize the importance of understanding it as an effective limit of a discrete interaction. By a classical model of discrete interaction we mean the replacement of the potential, representative of the continuous interaction, by the exchange of, in an evident abuse of language, classical 'quanta', little bits of well-defined amounts of energy-momentum. This classical quantum is emitted/absorbed in an instantaneous process caused also by the absorption/emission of a previous quantum. The emitted quantum travels towards its absorber-to-be which is causally reached after a

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time interval Δt . Its instantaneous absorption causes an immediate change in the energymomentum of the absorber and the immediate emission of a new quantum back to the first interaction body. All bodies, as well as their travelling quanta, propagate as free objects in the time interval between consecutive interactions (emission/absorption of quanta). Then, there is no potential energy, only kinetic energy. It is hoped that the discrete-to-continuous transition on such a naive classical model can mimic and enlighten some of the basic aspects of the quantum-to-classical transition on the description of nature. The search for a better understanding of quantum mechanics, which started with Einstein, de Broglie, Schrödinger and many other pioneers, including Bohm [1] with his well-known proposal, still goes on today with great intensity. Citing 't Hooft who also uses simple classical models with discrete time [2] in his search for a better understanding of quantum physics—'The dividing line between quantum physics and classical physics is more subtle than usually advertised' [3]. The wisdom of Khrennikov and Volovich [4] is that the continuous Newton model is just an approximation of physical reality and that most of the contradiction between the quantum formalism and Newtonian mechanics is, therefore, a consequence of this continuous approximation. This contribution goes along this line of thought.

A non-relativistic regime, which is appropriate for dealing with harmonic oscillator, is assumed. So, we will consider the relative movement of the reduced mass *m* of a non-relativistic two-body system. At the initial time t_0 it has the initial position \vec{r}_0 and momentum \vec{p}_0 ; it freely propagates on a rectilinear trajectory until its first subsequent interaction at

$$t_1 = t_0 + \Delta t_0$$
 $\vec{r}_1 = \vec{r}_0 + \Delta \vec{r}_0 = \vec{r}_0 + \frac{\vec{p}_0}{m} \Delta t_0$

that changes its momentum to

$$\vec{p}_1 = \vec{p}_0 + \Delta \vec{p}_0$$

It then follows, again, a rectilinear trajectory until its second interaction at

$$t_2 = t_1 + \Delta t_1$$

which changes its momentum to

$$\vec{p}_2 = \vec{p}_1 + \Delta \vec{p}_1.$$

So, at the *n*th interaction,

$$t_n = t_0 + \sum_{j=0}^{n-1} \Delta t_j$$
 (1)

$$\vec{p}_n = \vec{p}_{n-1} + \Delta \vec{p}_{n-1} = \vec{p}_0 + \sum_{i=0}^{n-1} \Delta \vec{p}_j$$
⁽²⁾

$$\vec{r}_n = \vec{r}_{n-1} + \Delta \vec{r}_{n-1} = \vec{r}_{n-1} + \frac{\vec{p}_{n-1}}{m} \Delta t_{n-1} = \vec{r}_0 + \frac{1}{m} \sum_{j=0}^{n-1} \vec{p}_j \Delta t_j.$$
(3)

The position vector, continually changing, describes a polygonal trajectory. There is a sudden change of momentum at each interaction point, a vertex of the polygon. Momentum is clearly a discrete parameter, or at least it changes discretely. The continuous variables time and position enter in the description of the motion as if they were also discrete parameters only because the interaction events are our reference points for counting time; this is not a lattice calculation. One cannot talk of force or acceleration, just of sudden change of momentum or velocity. They become effective concepts, valid only in appropriate limits, which are our objects of discussion here.

$$\Delta \vec{p}_{n-1} = -m\omega^2 \vec{r}_{n-1} \Delta t_{n-1} \tag{4}$$

which defines a discrete harmonic oscillator. Among numerously infinite possibilities we will consider the mathematically simplest one

$$\Delta t_n \equiv \alpha \tag{5}$$

$$\Delta \vec{p}_n \equiv -\alpha m \omega^2 \vec{r}_n \tag{6}$$

where α is a positive constant. Equations (2) and (3) then become

$$\vec{p}_n = \vec{p}_0 - m\alpha\omega^2 \sum_{j=0}^{n-1} \vec{r}_j$$
(7)

$$\vec{r}_n = \vec{r}_0 + \frac{\alpha}{m} \sum_{j_1=0}^{n-1} \vec{p}_j.$$
(8)

The recursive combination of equations (8) and (7) leads to

$$\vec{r}_n = \sum_{s=0}^{[n/2]} (\alpha \omega)^{2s} \vec{r}_n^{(s)}$$
(9)

and

$$\vec{p}_n = \sum_{s=0}^{[n/2]} (\alpha \omega)^{2s} \vec{p}_n^{(s)}$$
(10)

where $\vec{r}_n^{(s)}$ and $\vec{p}_n^{(s)}$ are polynomial functions of \vec{r}_0 and \vec{p}_0 , [n/2] is the largest integer in n/2, with

$$\vec{r}_n^{(0)} = \vec{r}_0 + \frac{\vec{p}_0}{m} \sum_{j_1=0}^{n-1} \alpha = \vec{r}_0 + \frac{\alpha \vec{p}_0}{m} n$$
(11)

$$\vec{p}_n^{(0)} = \vec{p}_0 - m\alpha \omega^2 \vec{r}_0 n \tag{12}$$

$$\vec{r}_n^{(s)} = -\sum_{j_1=0}^{n-1} \sum_{j_2=0}^{j_1-1} \vec{r}_{j_2}^{(s-1)}$$
(13)

$$\vec{p}_n^{(s)} = -\sum_{j_1=0}^{n-1} \sum_{j_2=0}^{j_1-1} \vec{p}_{j_2}^{(s-1)}.$$
(14)

 $\vec{r}_n^{(0)}$ would be the position at t_n if there were no interaction between t_0 and t_n , while $\vec{p}_n^{(0)}$ would be the momentum at t_n if all *n* interactions were equal to the first one; the vector $\vec{r}_n^{(1)}$ would be the final position if there were just one interaction in this time interval, and so on. Successive re-iterations on equations (13) and (14) lead to

$$\vec{r}_{n}^{(s)} = (-1)^{s} \sum_{j_{1}=0}^{n-1} \sum_{j_{2}=0}^{j_{1}-1} \dots \sum_{j_{2s}=0}^{j_{2s-1}-1} \vec{r}_{j_{2s-1}}^{(0)}$$
$$= (-1)^{s} \sum_{j_{1}=0}^{n-1} \sum_{j_{2}=0}^{j_{1}-1} \dots \sum_{j_{2s}=0}^{j_{2s-1}-1} (\vec{r}_{0} + \frac{\alpha \vec{p}_{0}}{m}n)$$
(15)

$$\vec{p}_{n}^{(s)} = (-1)^{s} \sum_{j_{1}=0}^{n-1} \sum_{j_{2}=0}^{j_{1}-1} \dots \sum_{j_{2s}=0}^{j_{2s-1}-1} \vec{p}_{j_{2s-1}}^{(0)}$$
$$= (-1)^{s} \sum_{j_{1}=0}^{n-1} \sum_{j_{2}=0}^{j_{1}-1} \dots \sum_{j_{2s}=0}^{j_{2s-1}-1} (\vec{p}_{0} - m\alpha\omega^{2}\vec{r}_{0}n).$$
(16)

Now, we use $n = \binom{n}{1}$ and the identity

$$\sum_{j=0}^{n-1} \binom{j}{k} = \binom{n}{k+1} \quad \text{for} \quad n > k \tag{17}$$

which can easily be proved by induction, to get

$$\vec{r}_n^{(s)} = (-1)^s \left(\vec{r}_0 \binom{n}{2s} + \frac{\alpha \vec{p}_0}{m} \binom{n}{2s+1} \right)$$
(18)

$$\vec{p}_n^{(s)} = (-1)^s \left(\vec{p}_0 \binom{n}{2s} - m\alpha \omega^2 \vec{r}_0 \binom{n}{2s+1} \right)$$
(19)

which, with equations (9) and (10), leads to

$$\vec{r}_n = \sum_{s=0}^{\lfloor n/2 \rfloor} (-1)^s \left[(\alpha \omega)^{2s} \vec{r}_0 {n \choose 2s} + (\alpha \omega)^{2s+1} \frac{\vec{p}_0}{m\omega} {n \choose 2s+1} \right]$$
(20)

$$\vec{p}_n = \sum_{s=0}^{[n/2]} (-1)^s \left[(\alpha \omega)^{2s} \vec{p}_0 \binom{n}{2s} - m\omega (\alpha \omega)^{2s+1} \vec{r}_0 \binom{n}{2s+1} \right].$$
(21)

For $n \gg 1$, however, the following approximation is valid

$$\binom{n}{k} \approx \frac{n^k}{k!} \tag{22}$$

and then

$$\sum_{s=0}^{\lfloor n/2 \rfloor} (-1)^s (\alpha \omega)^{2s} \binom{n}{2s} \approx \sum_{s=0}^{\lfloor n/2 \rfloor} (-1)^s (\alpha \omega)^{2s} \frac{n^{2s}}{(2s)!} \approx \cos(\alpha \omega n)$$
(23)

$$\sum_{s=0}^{[n/2]} (-1)^s (\alpha \omega)^{2s+1} \binom{n}{2s+1} \approx \sum_{s=0}^{[n/2]} (-1)^s (\alpha \omega)^{2s+1} \frac{n^{2s+1}}{(2s+1)!} \approx \sin(\alpha \omega n).$$
(24)

The middle terms of equations (23) and (24) are partial sums of the series representing the respective trigonometric functions. Therefore, these trigonometric functions can be seen as the asymptotic limits when $n \to \infty$ of the respective finite series of combinatorials. Then equations (20) and (21) become the general solutions of the standard (continuous interaction) harmonic oscillator, valid in this limit as an effective description. For large enough values of the number *n* of interaction events, the discrete changes of the parameters t_n , \vec{r}_n and \vec{p}_n are so small that they can effectively be described as if they were continuously changing under the action of a continuous interaction field, the harmonic potential. In this limit

$$H_n \equiv \frac{1}{2m}\vec{p}_n^2 + \frac{\alpha\omega^2}{2}\vec{r}_n^2 = H_0.$$
 (25)

Similar or related behaviour has been proved for classical electrodynamics [5], Newtonian gravitation [6], general relativity [7] and, generically, for field theory [8].

The harmonic potential appears as an effective continuous interaction with two asymptotic (in the limit of very large n) conditions:

$$n \gg 1$$
 (26)

and

$$\alpha \omega n = \omega (t_n - t_0) \ll 1 \tag{27}$$

for the trigonometric function retrieval. We made use of $t_n = t_0 + \alpha n$. In such a limit we can safely treat changes due to $\Delta n = 1 \ll n$ as time derivatives

$$\frac{\mathrm{d}}{\mathrm{d}t_n} = \frac{\mathrm{d}n}{\mathrm{d}t_n}\frac{\mathrm{d}}{\mathrm{d}n} = \frac{1}{\alpha}\frac{\mathrm{d}}{\mathrm{d}n}.$$
(28)

For keeping track of the order of what is being neglected with approximation (22) which leads to the continuous description, we use $\delta^{(k)}$ to indicate the *k*th largest contribution in a given neglected term,

$$\delta^{(1)}\binom{n}{k} = -\frac{n^{k-1}}{k!}\binom{k}{2} = -\frac{1}{2}\frac{n^{k-1}}{(k-2)!}.$$
(29)

Then, from equations (20) and (21)

$$\delta^{(1)}\vec{r}_n = \sum_{s=0}^{[n/2]} (-1)^s \left[(\alpha\omega)^{2s} \vec{r}_0 \delta^{(1)} \binom{n}{2s} + (\alpha\omega)^{2s+1} \frac{\vec{p}_0}{m\omega} \delta^{(1)} \binom{n}{2s+1} \right]$$
(30)

$$\delta^{(1)}\vec{p}_n = \sum_{s=0}^{[n/2]} (-1)^s \left[(\alpha\omega)^{2s} \vec{p}_0 \delta^{(1)} \binom{n}{2s} - m\omega(\alpha\omega)^{2s+1} \vec{r}_0 \delta^{(1)} \binom{n}{2s+1} \right]$$
(31)

from which we get

$$\delta^{(1)}\vec{r}_n = -(\alpha\omega)^2 \frac{n}{2}\vec{r}_n \qquad \delta^{(1)}\vec{p}_n = -(\alpha\omega)^2 \frac{n}{2}\vec{p}_n \tag{32}$$

so that

$$\delta^{(1)}H_n \approx H_n \left(\frac{\vec{p}_n \delta^{(1)} \vec{p}_n}{m} + m\omega^2 \vec{r}_n \delta^{(1)} \vec{r}_n\right) = -(\alpha \omega)^2 n H_n.$$
(33)

The fractional neglected terms, according to equation (27), are of the order of $\frac{1}{n} \ll 1$. Whereas the effective continuous limit is not affected by the choice made in equation (5), the neglected contributions $\delta^{(1)}$ are strongly affected. It is a characteristic of discrete interaction models that first-order corrections ($\delta^{(1)}$) to the effective continuous fields are proportional to n (or t_n). This may be not very relevant for the harmonic oscillator or any system with periodic motion, in contradistinction to, for example, a nearly radial escaping motion³ in a gravitational field⁴. In this case, a large/small n (or t_n) implies a large/small r_n . There are then three distinct scales to be considered:

1. *n* not large enough for the validity of the continuous interaction approximation. Necessarily, a discrete interaction description must be used. This corresponds to microscopic distances, a domain still lacking experimental data on gravity [10].

³ The anomalous acceleration observed in the Pioneer effect [9], not detected in planetary motion, is an example.
⁴ Shield effects prevent this happening with electromagnetic interactions, the other long-range fundamental interaction.

- 2. Very large *n* that validates the effective continuous field description, but not large enough so that the $\delta^{(1)}$ -corrections may be neglected. This corresponds to the solar system size scale where the continuous (Newton and Einstein) descriptions have solid experimental confirmation [11].
- 3. Values of *n* so large that the $\delta^{(1)}$ -corrections cannot be neglected. It is well known that for galactic and cosmological distance scales the standard continuous descriptions of gravity require the *ad hoc* assumption of, up-to-now, unseen new forms of (dark) matter and energy [12]. These new actors are called for the compensation of the inability of the standard field to produce a repulsive interaction component (this *wrong*-sign interaction is necessary for fitting the data) from their known sources [13]. The $\delta^{(1)}$ -correction terms can do that [6] and provide a natural interpretation for this interaction *wrong*-sign. This, however, is in direct contradiction to the predictions of quantum gravity for which the effects of discrete gravity are to be revealed only at ultra-high energies or ultra-short distances.

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References

- Bohm D and Hiley B 1993 The Undivided Universe: An Ontological Interpretation of Quantum Mechanics (London: Routledge and Kegan Paul)
- [2] 't Hooft G 2001 Quantum mechanics and determinism *Preprint* hep-th/0105105
- [3] 't Hooft G 2001 How does God play dice? (Pre)-determinism Preprint hep-th/0104219
- [4] Khrennikov A Yu and Volovich Ya I 2002 Interference effect for probability distributions of deterministic particles and discrete time leads to quantum-like interference of deterministic particles *Proc. Int. Conf. Quantum Theory: Reconsideration of Foundations (Ser. Math. Modelling in Phys., Engin., and Cogn. Sc.)* (Växjö, Sweden: Växjö University Press) pp 455–62 and 441–54 (*Preprints* quant-ph/0111159 and quant-ph/0203009)
- [5] de Souza M M 1997 Classical electrodynamics and the quantum nature of light J. Phys. A: Math. Gen. 30 6565–85 (Preprint hep-th/9610028)
 - de Souza M M 2002 Dynamics and causality constraints *Braz. J. Phys.* **32** 609–16 (*Preprint* hep-th/0006214) de Souza M M 1999 Discrete gauge fields *Preprint* hep-th/9911233
- [6] de Souza M M 2001 Gravity and antigravity with discrete interactions Preprint gr-qc/0111052
- [7] de Souza M M and Silveira R N 1999 Discrete fields and general relativity Class. Quantum Grav. 16 619 (Preprint gr-qc/9810040)
 - de Souza M M 2001 Discrete fields, general relativity and other possible experimental evidences *Preprint* gr-qc/0112070
- [8] de Souza M M 2000 Conformally symmetric massive discrete fields Preprint hep-th/0006237
- [9] Anderson J D, Laing P A, Lau E L, Liu A S, Nieto M M and Turyshev S G 2002 Search for a standard explanation of the pioneer anomaly *Mod. Phys. Lett.* A 17 875–86 (*Preprint* gr-qc/0107022,0205059,9808081,9903024)
- [10] Antoniadis I, Arkani-Hamed N, Dimopoulos S and Dvali G R 1998 Nucl. Phys. B 516 70–82 (Preprint hep-ph/9710204)
 - Ferrara S, Kounnas C and Zwirner F 1994 Nucl. Phys. B 429 589
- [11] Will C M 2001 The confrontation between general relativity and experiment Liv. Rev. Rel. (Preprint gr-qc/0103036)
- [12] Perlmutter S et al 1998 Nature 391 51
 Zwicky F 1933 Helv. Phys. Acta 6 110
 Sellwood J A and Kosokowsky A 2000 Does dark matter exist? Preprint astro-ph/0009074
- [13] Manheim P D 2000 Attractive and repulsive gravity Found. Phys. 308 709-46 (Preprint gr-qc/0001011)